

ELIPSE

Obtencion de posicion y velocidad a aprtir de elementos orbitales.
Curso Mecanica Celeste 2005. Tabare Gallardo.

unidad de tiempo = dia
unidad de distancia = ua
unidad de masa = sol

```
M := 1.0
m := 0.0009548
kgauss := 0.01720209895
μ := kgauss * kgauss * (M + m)
a := 5.179995814935037
e := 0.04871959170445102
i := 1.3046833201072958 * Pi / 180
ω := -84.90588389067327 * Pi / 180
Ω := 100.48313071118518 * Pi / 180
T := -81.299453540603
t := 0.0
```

- Esta kepler me la paso Daniel Chagas.

```
kepler[tt_, ε_, mm_] :=
Module[{lista = {}, ee = tt}, For[j = 1, j ≤ mm, j++,
ee = tt + ε * Sin[ee]; lista = Append[lista, ee]]; Return[lista]]
```

$$n := \sqrt{\frac{\mu}{a^3}}$$

```
p := a * (1 - e * e)
h := √(p * μ)
```

```
AM1 = n * (t - T)
```

```
0.118681
```

```
AM = Mod[AM1, 2 * Pi]
```

```
0.118681
```

```
kepler[AM, e, 20]
```

```
{0.12445, 0.124729, 0.124742, 0.124743, 0.124743,  
 0.124743, 0.124743, 0.124743, 0.124743, 0.124743,  
 0.124743, 0.124743, 0.124743, 0.124743, 0.124743,  
 0.124743, 0.124743, 0.124743, 0.124743, 0.124743}
```

```
AE = Last[%]
```

```
0.124743
```

```
r = a * (1 - e * Cos[AE])
```

```
4.92959
```

```
 $\theta := \text{Sign}[AE] * \text{ArcCos}\left[\frac{p/r - 1}{e}\right]$ 
```

```
{r,  $\theta$ }
```

```
{4.92959, 0.130959}
```

```
 $\xi := r * \text{Cos}[\theta]$ 
```

```
 $\eta := r * \text{Sin}[\theta]$ 
```

```
 $\xi := 0$ 
```

```
vecpos := { $\xi$ ,  $\eta$ ,  $\xi$ }
```

```

xp := -h / p * Sin[θ]
yp := h / p * (e + Cos[θ])
zp := 0

```

```

vecvel := {xp, yp, zp}

```

```

P1={{Cos[ω],-Sin[ω],0},{Sin[ω],Cos[ω],0},{0,0,1}};

```

```

P2={{1,0,0},{0,Cos[i],-Sin[i]},{0,Sin[i],Cos[i]}};

```

```

P3={{Cos[Ω],-Sin[Ω],0},{Sin[Ω],Cos[Ω],0},{0,0,1}};

```

```

P=P3.P2.P1;

```

```

veclipos = P.vecpos

{4.53377, 1.9323, -0.10954}

```

```

veclivel = P.vecvel

{-0.00306656, 0.00732014, 0.0000383417}

```

```

veclivel * 365.25

{-1.12006, 2.67368, 0.0140043}

```